# Telescopes

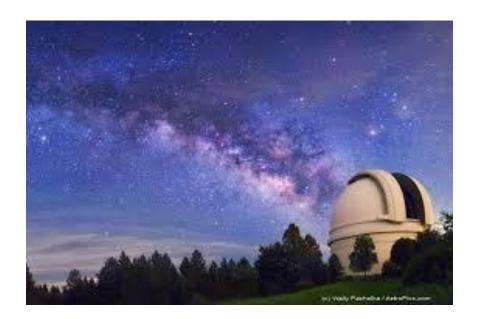


Astronomy

### Introduction

#### Telescopes

- Section 1: Telescopes
- Section 2: Resolving Power
- Section 3: Detecting Light
- Section 4: Telescopes on the Ground and in Space
- Section 5: Observatories



#### Introduction

- A \_\_\_\_\_ enables an astronomer to observe things not visible to the naked \_\_\_\_\_
  - Human eye sight can't see:
    - \_\_\_\_\_ objects
    - Fine details
    - Long distances
  - We rely off of light input in order to see images
    - No light = no \_\_\_\_\_
- Telescopes increase the image \_\_\_\_\_ for distances and adds light to help us see

#### Light-Gathering Power

- ▶ In order for us to see, \_\_\_\_\_ emitted or \_\_\_\_\_ from it need to strike the retina of the eye
- Photon particle of \_\_\_\_\_
- We can only see based on the number of photons coming in
- Some telescopes "collect" and "\_\_\_\_\_" photons in to our eye\_\_\_\_\_

Cornea

#### Light-Gathering Power

 Light - Gathering Power - a large \_\_\_\_\_ mirror used in telescopes to help collect and funnel

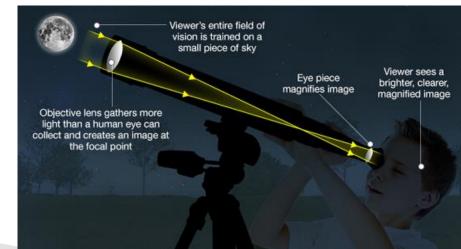
These give us brighter \_\_\_\_\_

The bigger the lens or mirror, the more photons can be caught  $(\pi r^2)$ 

An increase from 4 to 6 m in diameter means a difference

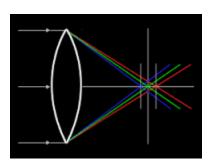
in size and capability

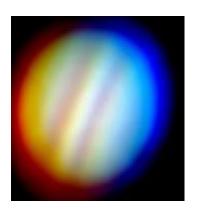
•  $2^2 = 4$  compared to  $3^2 = 9$ 



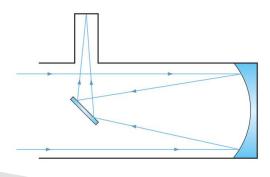
- Once light has been gathered, it needs to be focused
- Refractors refracting telescopes where \_\_\_\_\_\_ is gathered and focused by a \_\_\_\_\_
- Refraction \_\_\_\_\_of light rays
  - This happens when light moves from one substance to
    - another
    - Ex: air to water

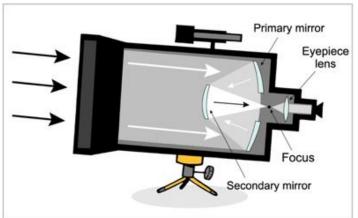
- Having lenses in large telescopes can have disadvantages:
  - Extremely \_\_\_\_\_ and tough to fabricate (make)
  - The lens has to be connected on the \_\_\_\_\_ and it causes it to \_\_\_\_\_ which distorts the images
  - Transparent materials bring light of different colors to focus at slightly different distances from the lens
    - Chromatic Aberration color flawed images





- Having lenses in large telescopes can have disadvantages (cont.):
  - Many lens materials completely \_\_\_\_\_ short wavelength light
- Most modern telescopes use \_\_\_\_\_ to help with the issues lenses cause
- Reflectors telescopes that use mirrors to \_\_\_\_\_ light



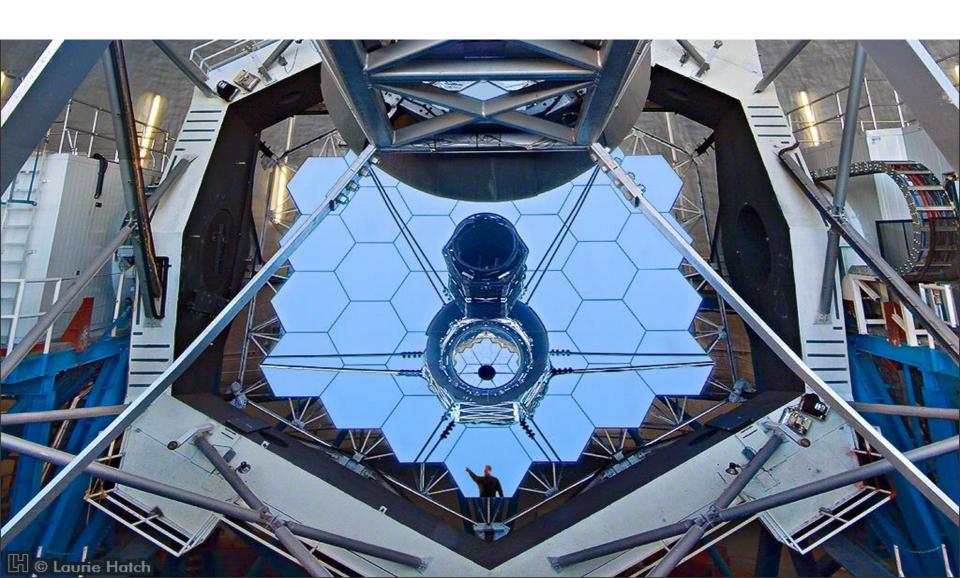


### Focusing Light

- ▶ The mirrors on \_\_\_\_\_ hare made of glass have been shaped to a smooth curve, polished, and then coated with a thin layer of \_\_\_\_\_ or some other highly reflective metal
- Because the light doesn't pass through the \_\_\_\_\_, the \_\_\_\_ don't get distorted
- Also, the mirror can be supported from the back and won't sag

Winner across the board!

**UCLA Galactic Center's Reflector** 



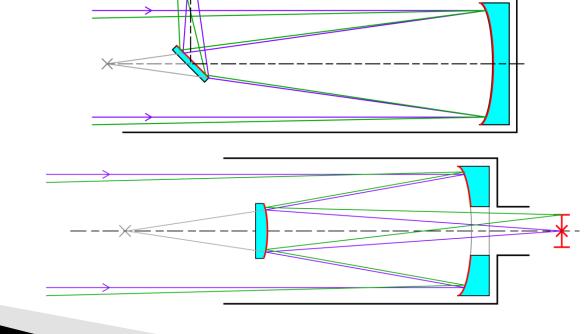
### **Focusing Light**

▶ In order for the mirror system to work, you need at least \_\_\_\_\_ of them

The primary mirror reflects the \_\_\_\_\_ light

The secondary mirror angles that light towards the focus

(eye piece)



- Most telescopes are mounted on big \_\_\_\_\_ that allow them to \_\_\_\_\_ on objects as they travel across the sky
  - trying to get the telescope to move smoothly with all of the mirrors in the correct locations to keep the sharp image is quite a challenge and requires a lot of \_\_\_\_\_\_
- When they move, it causes the glass in the mirror to \_\_\_\_\_, especially in really big reflectors
- A lot of the current reflectors are using multiple \_\_\_\_\_ mirrors instead of one large one to fix that problem

### **Focusing Light**

Lasers are used to keep the mirrors \_\_\_\_\_\_ properly as the reflector moves



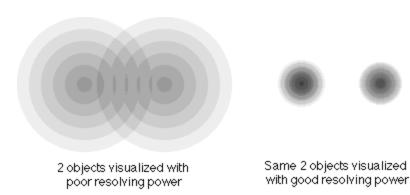
#### Introduction

- Stars and other objects that lie very close together (land formations on planets, actual planets from far away, etc.) may be two \_\_\_\_\_ objects but can be visualized through our eyes as one
  - They'll \_\_\_\_\_ or mesh together if small enough or far enough away



#### **Resolving Power**

- Resolving Power a telescope's ability to \_\_\_\_\_
  objects or details on those objects
- Limited by the wave nature of \_\_\_\_\_
  - Ex: two stars that are only separated by a small angle (from our view) may have light waves that cross (this is why we see them as one), so for the scope to distinguish them, the light waves have to be careful not to cross completely and blend together



#### **Resolving Power**

- Diffraction situation where small, \_\_\_\_\_, light waves are produced when bigger ones pass through an \_\_\_\_\_
- 1 source of light becomes surrounded by rings of secondary light that can even sort by color wavelength

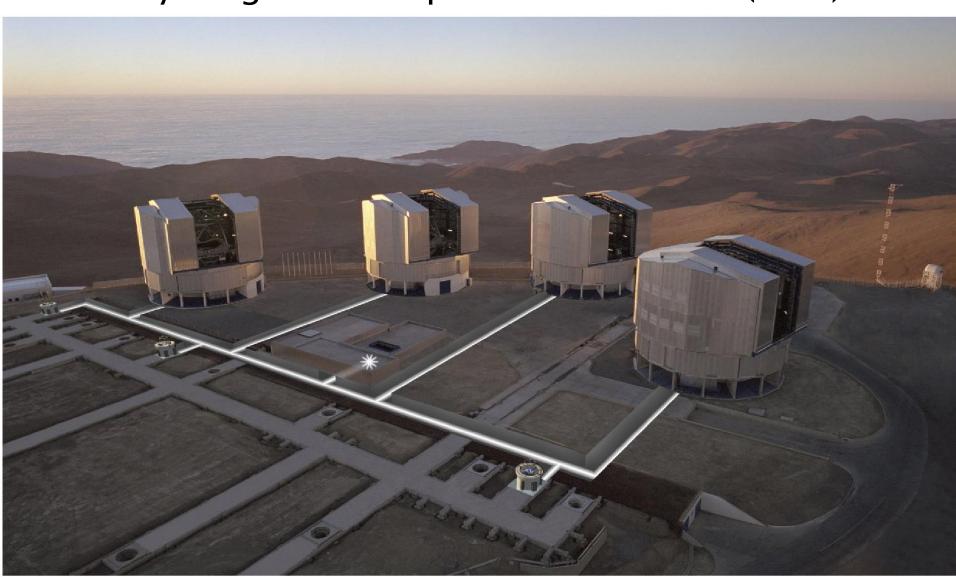




### **Resolving Power**

- Diffraction effects \_\_\_\_\_ be completely eliminated, but they can be reduced
  - Enlarge the opening through which the light passes so that the waves don't mix as much
- Interferometer two or more widely spaced
   \_\_\_\_\_ that direct the light to a common
   detector that combines the separate light \_\_\_\_\_\_

The Very Large Telescope Interferometer (VLTI)

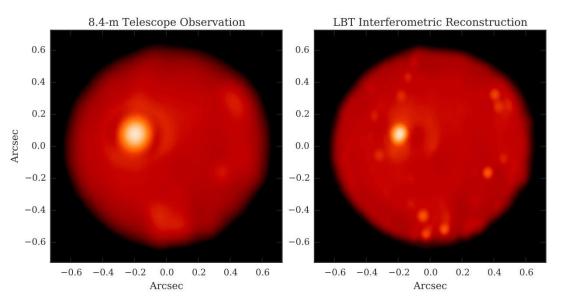


#### **Resolving Power**

The interferometer is named for it ability to separate waves that "\_\_\_\_\_\_" with each other

Important to remember: the strength of the telescope is one thing, but the ability to \_\_\_\_\_\_, through the use of a interferometer, can be more

valuable



- Once light gets collected, it must be \_\_\_\_\_ and recorded
  - Old days: an astronomer detected the light that was collected through the eye piece of the telescope and recorded the light by \_\_\_\_\_ the image that was seen through the scope
- Many celestial objects are too small and too faint for us to see with our own eyes
  - At most, many of the celestial objects emit just a few photons of light by the time they reach us

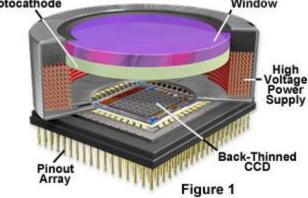


- If you were to look at a near \_\_\_\_\_ through a large observatory scope, it may take several hours of light capture to assemble a picture
- ▶ To see faint objects, astronomers use different kinds of \_\_\_\_\_ that are able to store light in some way to create images
  - Can be done both chemically with photos and digitally with detectors like ones in video camcorders

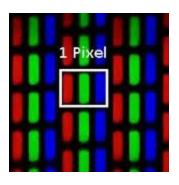


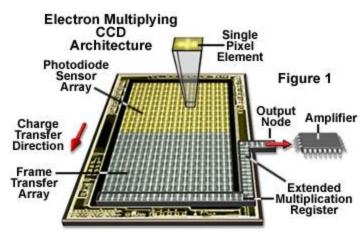
- From the late 1800s to the 1980s, astronomers usually used \_\_\_\_\_ film to record the light from the bodies they were studying
  - Film absorbs photons that cause a chemical change, making the film dark where the light hit and thus creating the photograph
  - Not very effective though... it took hours to capture and even longer to develop
- Today: almost all \_\_\_\_\_ detectors

- Charge-Coupled Device (CCD) \_\_\_\_\_ detector that can make pictures virtually \_\_\_\_\_ from photographs in their detail and with a sensitivity to faint light
  - Approximately \_\_\_\_\_ greater
  - The light coming in strikes a semiconductor surface which allows electrons to move within the material
  - The surface is divided into a bunch of small squares called pixels where the electrons are stored



- The number of electrons in each \_\_\_\_\_ is proportional to the number of photons hitting it
- The device is then connected to computer that scans the detector, counting the number of electrons in each pixel and generating a picture
  - Similar to how \_\_\_\_\_ pixels are small dots all put together to create the image you see on the screen





### Visible Light

- CCDs are extremely efficient and can record \_\_\_\_\_ of the photons striking them, allowing astronomers to record images much faster than with film
- Because they are \_\_\_\_\_ images, they can be altered by sharpening them, removing extraneous light, and enhancing contrast to help produce a

great image

- The visible light we see is a \_\_\_\_\_ portion of the full electromagnetic spectrum
- Many \_\_\_\_\_ objects give off wavelengths we can't see because it doesn't fall within the visible light spectrum
  - Ex: cold gas clouds emit (give off) radio waves so we use radio detectors to compile those images



- - These can also be aligned into \_\_\_\_\_ that cross over entire continents!
    - This is because the radio waves are so long





- Many different celestial objects radiate \_\_\_\_\_\_
  energy
- ▶ This is tricky...
  - Astronomers can use infrared telescopes, but the telescopes themselves have the ability to give off \_\_\_\_\_ infrared energy
  - These scopes have to be in really low temperatures and away from any \_\_\_\_\_ (walls) to allow that infrared to collect around it

- X rays are even tougher...
  - X rays are easily absorbed by \_\_\_\_\_ if they hit directly
  - If they hit at a \_\_\_\_\_\_, horizontal angle, they can be reflected
    - Very similar to how a rock can skip on water
- X ray scopes are curved \_\_\_\_\_ that gradually direct the photons towards the detector

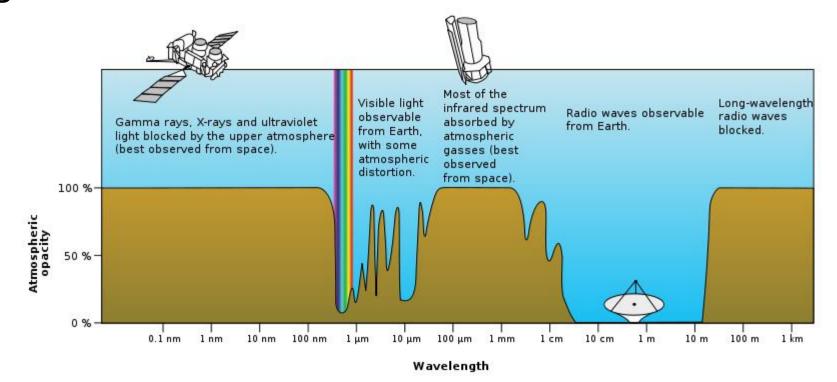


- Because we can't see x rays and \_\_\_\_\_, we use false \_\_\_\_\_ images to form images with these
- The colors represent different amounts of radiation
  It's translated to "color"
- Most of the \_\_\_\_\_ we can't see have a hard time getting through Earth's atmosphere
- Best option: put them up in \_\_\_\_\_!



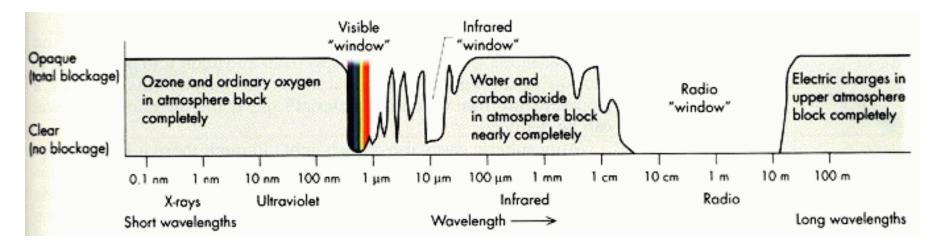
#### Introduction

 Atmospheric Window – the \_\_\_\_\_ region that gives one the ability to peer out into space from the ground



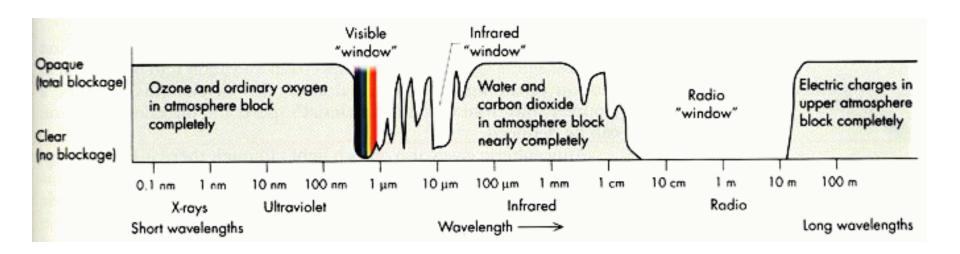
#### Introduction

- \_\_\_\_\_ in our atmosphere absorb infrared, UV, and shorter wavelengths
  - Ozone
  - Carbon dioxide
  - Water



#### Introduction

Because the \_\_\_\_\_ can absorb so many different wavelengths, it is important for satellites and \_\_\_\_\_ to get out past the atmosphere and go to space



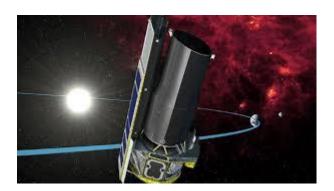
#### Introduction

Some of the scopes out there designed to detect

these absorbed \_\_\_\_\_:

- Space Telescope (HST)
- Extreme \_\_\_\_\_ Explorer (EUVE)
- Spitzer \_\_\_\_\_ Space Telescope
- \_\_\_\_\_\_ X ray Telescope Satellite







#### Introduction

- Some of the scopes could be launched and will work for very long periods of time, others will fall short
  - Ex: Spritzer \_\_\_\_\_
    - The Spritzer scope needs liquid helium to keep it cool enough to detect infrared
    - Once the liquid \_\_\_\_\_ ran out, it can no longer detect those waves
    - It collects minimal information in regards to gamma and Xray, but that's it

#### **Atmospheric Blurring**

- Most popular scope: Hubble Space Telescope (HST)
  - Can observe at \_\_\_\_\_ wavelengths and portions of infrared and \_\_\_\_\_
  - Contains a primary mirror that's \_\_\_\_\_ meters in diameter
  - Smaller than most land scopes
  - Puts out fully detailed images because it can dodge the "blurring" of the atmospheric components
    - \_\_\_\_\_
    - Dust
    - pollution

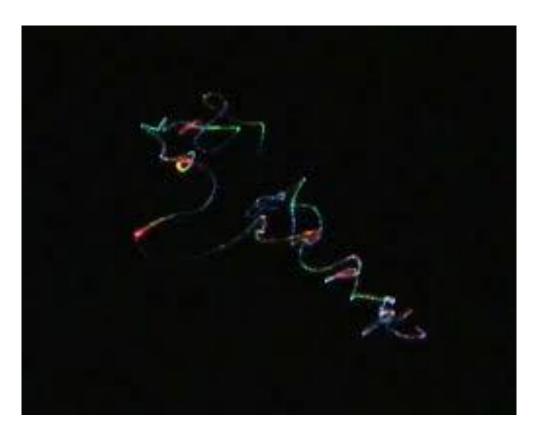


### **Atmospheric Blurring**

- Stars have always been known to "\_\_\_\_\_\_\_'
- Scintillation a situation where atmospheric \_\_\_\_\_ refracts (bends) the star's light
  - Atmospheric variations can be in density caused by small temperature changes
  - Because of this, the pathways of light can change direction and cause \_\_\_\_\_ to occur with light from other directions
  - Ex: looking at a penny under moving water
    - The penny looks like it's dancing when in reality it's the light bending and causing that view to be recorded by you

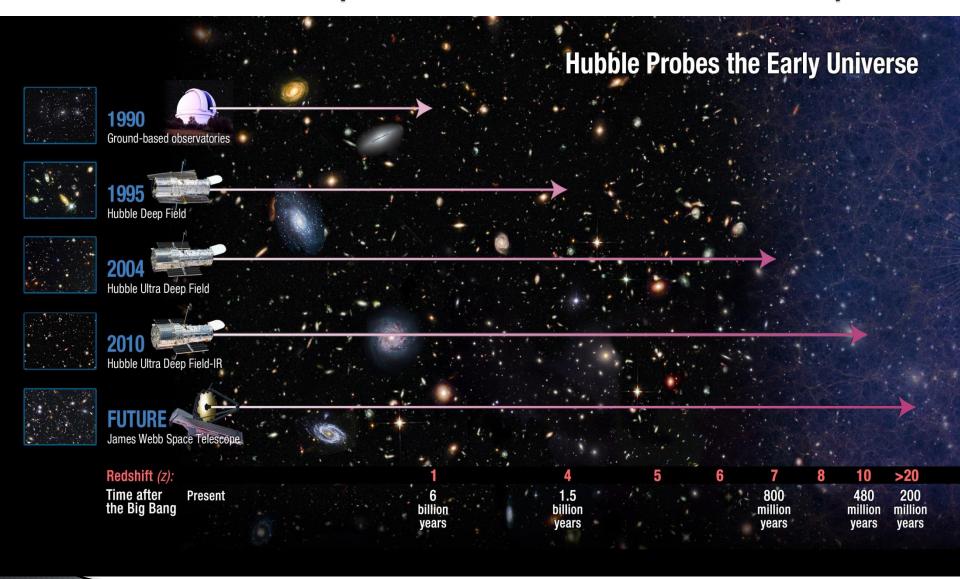
## **Atmospheric Blurring**

- Scintillation of Sirius Over a Period of Time
  - Notice the color expansion and \_\_\_\_\_ light pattern



### **Atmospheric Blurring**

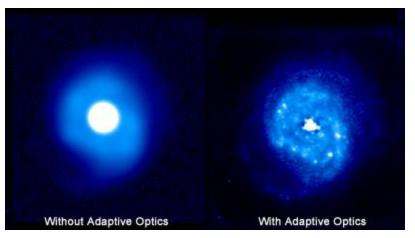
- Seeing the \_\_\_\_\_ of light by means of the atmospheric changes
- This gets \_\_\_\_\_ once you get past the atmosphere
- Continued repair to the Hubble allows it to stay as the #1 viewing tool in space
  - The launch of the \_\_\_\_\_ Telescope will replace it (fingers crossed!)



- There are ways to go about \_\_\_\_\_ the distortion from the atmosphere if on the ground
- Learning the object being studied to a known star
  - knowing exactly how the known star is being distorted, helps astronomers compare those same changes to the object being studied
  - Hard part: it's tough to get a good \_\_\_\_\_ star to be close to the object being studied
  - Fix: use a \_\_\_\_\_!



- A powerful laser \_\_\_\_\_ is projected into the sky to create an \_\_\_\_\_ star for comparison
  - the distortions of the artificial star image are recorded by a computer and adjusted by actuated mirrors
- Adaptive Optics using \_\_\_\_\_ (moving motors)
  on correcting mirrors that cancel out the
  distortions created by the \_\_\_\_\_\_



- These types of techniques give us great \_\_\_\_\_ into without the hassle and expenses of \_\_\_\_\_ into space
- They also give us the opportunity to construct large scale \_\_\_\_\_, much greater in size than ones that would need to go to space
- \_\_\_\_\_ is still very important!
  - Good spots: American Southwest, Australia, Hawaii, etc.



- Unforeseen obstacle: \_\_\_\_\_ pollution
  - Light pollution makes it really difficult to be able to see
  - Most of this light pollution leaks from \_\_\_\_\_ areas
  - With increases in world wide population and technology, light pollution is getting exponentially higher



#### Introduction

- The major telescopes used by astronomers are extremely \_\_\_\_\_ to manufacture
- Some of the largest ones are national or \_\_\_\_\_\_ facilities because of their expense
  - Many colleges and universities have their own observatories for research and instruction
- Some large \_\_\_\_\_ groups have them, too
  - Ex: Carnegie Institution



#### Introduction

 There are observatory telescopes on every continent, including \_\_\_\_\_

The cold and extremely dry air gives leeway to great

shots of the sky



### Introduction

- Twin \_\_\_\_\_ Telescopes
  - Two 10-meter scopes
  - Can operate individually or as an \_\_\_\_\_
  - Located in Hawaii



### Introduction

- Large Binocular Telescope
  - Located in \_\_\_\_\_
  - 8.4 m each
  - Even though it's one major unit, it can be used as an

\_\_\_\_\_, too



### Introduction

- Very Large Telescope (VLT)
  - \_\_\_\_\_ telescopes
  - 8.2 meter scopes
  - Work together as an interferometer
  - Located in \_\_\_\_\_

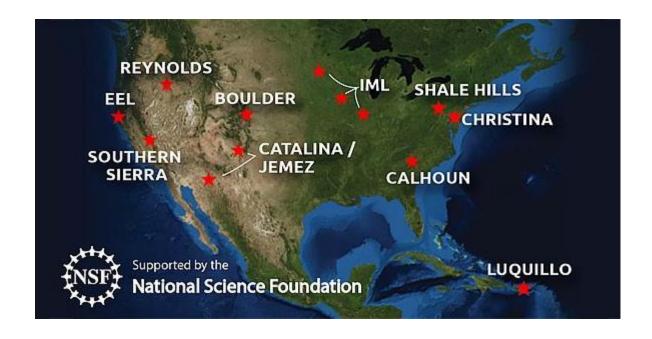


#### Introduction

With the high expense \_\_\_\_\_, the ability to study the sky is a \_\_\_\_\_ effort

Observatories from all over the globe share their
 \_\_\_\_\_ in order to help advance the field of

astronomy



## Going Observing

- In order for an astronomer to use the observatory scope, a \_\_\_\_\_ has to be submitted explaining what he/she wants to focus on and \_\_\_\_\_
- They also have to show that the scope's \_\_\_\_\_ and capabilities are what's needed to study the object
- Proposals are screened by a \_\_\_\_\_ that then allocate telescope time according to the scientific merits of the proposals

On the summit of Mauna Kea, Island of Hawai'i

## Going Observing

- The runs that get \_\_\_\_\_ are set up for several \_\_\_\_ in a row and are subject to whatever conditions may be present
- Some runs get great results, some are at a total loss due to storms or other inclement weather
- Mostly all scopes are run by \_\_\_\_\_ and software, allowing the observer to control the scope from an observation room or even from another location - most \_\_\_\_\_



### Computers

- Being able to operate a \_\_\_\_\_ and program software are more valuable \_\_\_\_\_ in astronomy than knowing how to use a telescope
- Computers are used to:
  - Solve \_\_\_\_\_
  - Move the telescope
  - Feed the information to detectors
  - Convert the data obtained by the \_\_\_\_\_
  - Communicate with other astronomers and observatories

