

## Introduction

#### Sections:

- Components of the Solar System
- Formation of Planetary Systems
- Other Planetary Systems



#### The Sun

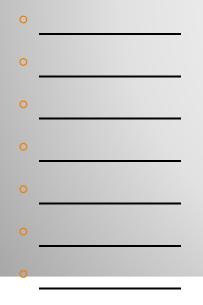
- The Sun is a star
- Star ball of \_\_\_\_\_ gas whose light and heat are generated by \_\_\_\_\_ reactions in its core
- It's the largest \_\_\_\_\_ in the solar system
  - More than 700x the mass of the other objects put together
  - Its gravitational force holds the other planets in place
- Solar System the \_\_\_\_\_ domination of the planets by the Sun

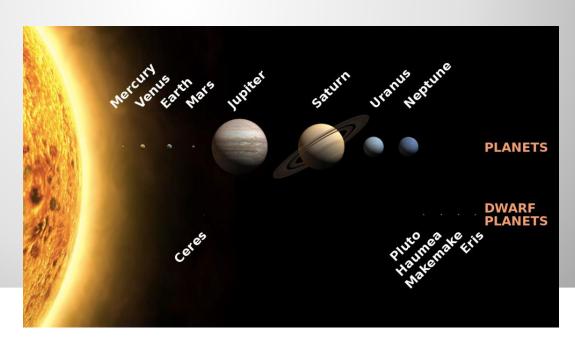
#### The Sun

- Mostly hydrogen and helium
  - About \_\_\_\_\_ % H
  - About \_\_\_\_\_ % He
- Contains small components of:
  - Carbon
  - Iron
  - \_\_\_\_
    - · All in a \_\_\_\_\_ form!
    - We can tell based off of the spectrum of light it emits

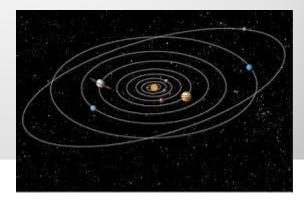


- Planets are much \_\_\_\_\_ than the Sun
- They emit no \_\_\_\_\_\_ light of their own
  - They do shine by reflecting the Sun's light
- Planets in order:





- The \_\_\_\_\_ of all of the planets around the Sun are mostly \_\_\_\_ and almost about the same plane, \_\_\_\_
  - It almost looks like a spinning pancake with the planets traveling around the Sun in the same direction
- The planets' rotation around the sun is

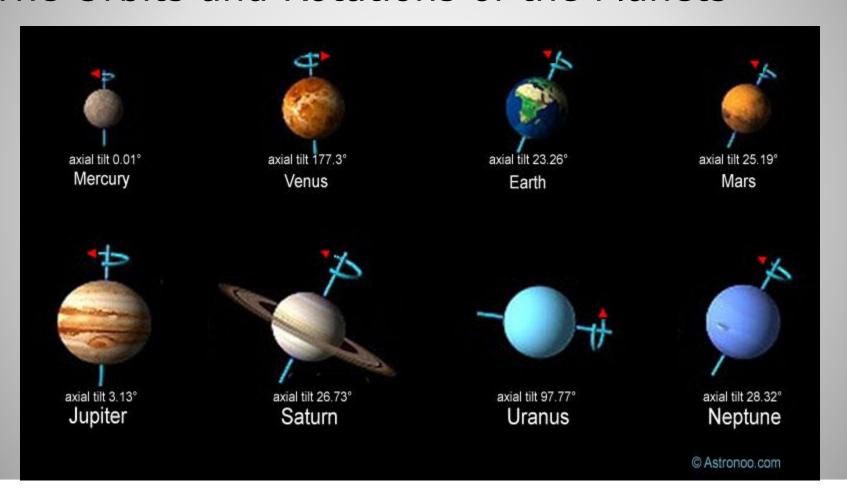


- As the planets orbit, each "\_\_\_\_\_ " on its rotation axis
  - The angle of the tilt has to do with how far off of the horizontal plane it is
- Generally, this \_\_\_\_\_ is in the same direction as the orbit around the Sun
- 2 exceptions to this:
  - 0
  - 0



- \_\_\_\_\_ has an extremely large tilt to its rotation axis (\_\_\_\_\_\_\_°)
- \_\_\_\_\_ rotation axis has such a large tilt that it actually spins backwards (\_\_\_\_\_°)
  - Still orbits in the same direction as the others around the Sun
- \_\_\_\_\_ rotation when a planet's rotation axis is so steep that it spins backwards





The Orbits and Rotations of the Planets

- These two \_\_\_\_\_ (same \_\_\_\_\_ orbit and flat \_\_\_\_\_ ) are the most fundamental features of the Solar System
- A third factor is that there are two different

types of planets:

- •
- 0
- Based on:
  - Size
  - 0
  - Location in the Solar System

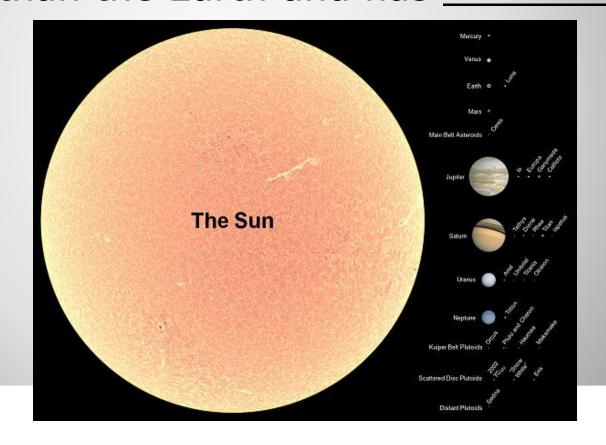
# Components of the Solar System Two Types of Planets Inner Planets – \_\_\_\_\_ bodies

- with relatively thin or no atmospheres
  - Mercury
  - 0
  - Earth
- 0
- Outer Planets \_\_\_\_\_ and liquid planets that are much \_\_\_\_\_ and have deep,
  - \_\_\_\_\_ atmospheres
  - Jupiter
  - 0
  - Uranus
  - 0

Two Types of Planets

 Jupiter is more than \_\_\_\_\_ larger in diameter than the Earth and has

its mass



- \_\_\_\_\_ and \_\_\_\_ are how we describe the planets
- Rock material composed of silicates
  - Silicates composed of \_\_\_\_\_\_, O, and other heavier elements like Al, Mg, S, and Fe
- Ice frozen liquids and gases
  - Such as:
    - Regular \_\_\_\_\_ (H<sub>2</sub>O)
    - Frozen \_\_\_\_\_ dioxide (CO<sub>2</sub>)
    - Frozen \_\_\_\_\_ (NH<sub>3</sub>)
    - Frozen \_\_\_\_\_ (CH<sub>4</sub>)



- Looking at the \_\_\_\_\_ solar system, rock is \_\_\_\_ because of the amount of hydrogen
  - Because of the heat near the sun, the carbon dioxide, methane, water, and ammonia can't condense to mingle with it
- The outer planets have no "\_\_\_\_\_\_\_\_"
  - Their \_\_\_\_\_ thicken with depth and eventually convert to liquid
  - Therefore we can't "\_\_\_\_\_ " on Jupiter or the other outer planets

- - Because they resemble \_\_\_\_\_
- \_\_\_\_\_ Planets Jupiter, \_\_\_\_\_ ,
   Uranus, and Neptune
  - Because they resemble

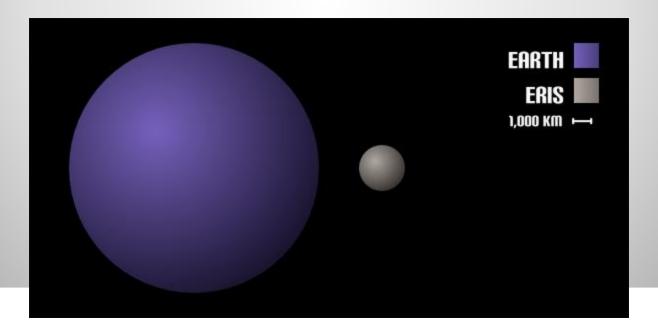


Components of the Solar System Two Types of Planets • Why no Made if and Pluto Odd Super \_\_\_\_\_ in comparison Astronomers found others similar to it Dwarf Planets – objects that orbit the , are massive enough that their compresses them into an approximately spherical shape, but have not swept their orbital region clear of other

objects that add up to a comparable mass

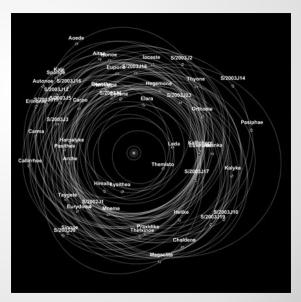
as the

- The discovery of \_\_\_\_\_ in \_\_\_\_ is what set the demotion of
  - Its closer than Pluto and also larger in size, but still fits the \_\_\_\_\_ planet criteria



#### Moons

- As the planets orbit the Sun, most are orbited by other \_\_\_\_\_
- Moons:
  - Jupiter:
  - Saturn:
  - Uranus:
  - Neptune: \_\_\_\_\_
  - Mars:
  - Earth: \_\_\_\_\_



- Even dwarf planets can have moons
  - Ex: Eris has \_\_\_\_\_

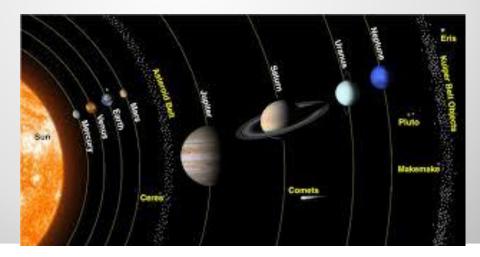
**Asteroids and Comets** 

- Asteroids and comets are far \_\_\_\_\_ than \_\_\_\_ objects
- Asteroids \_\_\_\_\_ or \_\_\_\_ objects
   with diameters that range from few meters
   up to about 1000 km
- Comets \_\_\_\_ objects about 10 km or less in diameter that grow huge tails of gas and dust as they near the Sun and are

partially \_\_\_\_\_ by heat

**Asteroids and Comets** 

- These two are not only different in \_\_\_\_\_
   but also their \_\_\_\_\_ in the solar system
- Asteroid Belt large \_\_\_\_\_ between the orbits of \_\_\_\_\_ and Jupiter where asteroids orbit the Sun



**Asteroids and Comets** 

- Most comets orbit far beyond Neptune
- Oort Cloud \_\_\_\_\_ region that completely \_\_\_\_ the solar system
  - Extends from about 40,000 to 100,000 AU from the Sun
  - Some can come closer
- Kuiper Belt the disk-like \_\_\_\_\_ of icy objects that lies just beyond the \_\_\_\_\_ of \_\_\_\_ of \_\_\_\_ of \_\_\_\_ and extends to be about 50 AU

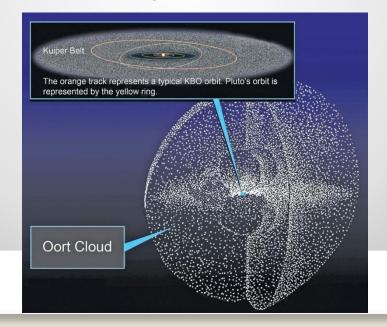
from the Sun

Composition – Inner and Outer Planets

- We can detect a planet's \_\_\_\_\_ a couple different ways
- Using its \_\_\_\_\_\_, we can measure its atmospheric composition and get some info about its \_\_\_\_\_\_ rocks (if they're there)
- We use \_\_\_\_\_ waves to tell us about Earth's interior and even though we haven't been able to do that with the other rock planets yet, it would tell us a lot of

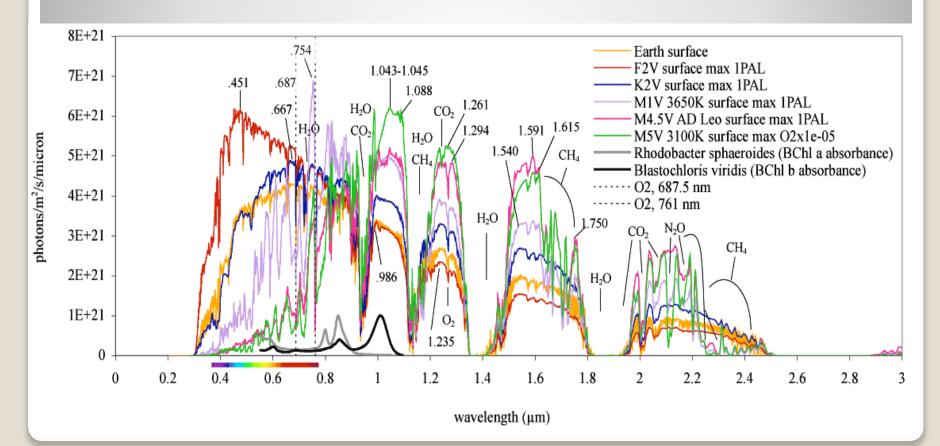
**Asteroids and Comets** 

- The \_\_\_\_\_ cloud and \_\_\_\_\_ put together can hold:
  - more than 1 \_\_\_\_\_ (1x10<sup>12</sup>) comets
  - thousands of \_\_\_\_\_ objects
  - Several dozen dwarf planets including \_\_\_\_\_



Composition – Inner and Outer Planets

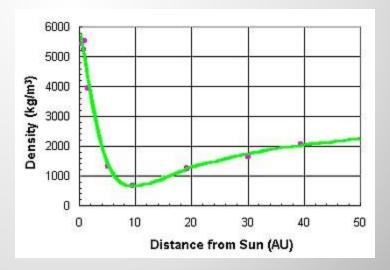
Earth – detected by \_\_\_\_\_ and simplified by



Composition – Inner and Outer Planets

 For \_\_\_\_\_ planets, we can't use any type of \_\_\_\_ work and the spectrum only takes care of the surface and atmosphere

The simplest \_\_\_\_\_ to use is planetary



- The average \_\_\_\_\_ of a planet is its mass divided by its \_\_\_\_\_
  - We can \_\_\_\_\_ a planet's mass by observing the orbital motion of one of its moons or a passing spacecraft
  - Then we can calculate the volume of the planet using one of Kepler's formulas from his third law

$$M =$$

**Planetary Density** 

- Volume
  - Variable:
    - *R* \_\_\_\_\_ of the planet
- Mass
  - Variables:
    - d \_\_\_\_\_ of object from the planet
    - *G* \_\_\_\_\_ constant
    - *P* \_\_\_\_\_ period
- Density

Μ	=

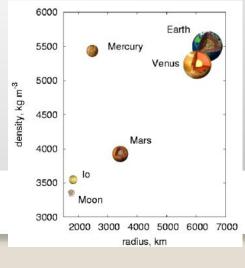
D =\_\_\_\_

Planetary Density

Once the planet's \_\_\_\_\_ density is known, we can compare it with the density of the \_\_\_\_\_, \_\_\_ materials to find what would \_\_\_\_\_ match up!

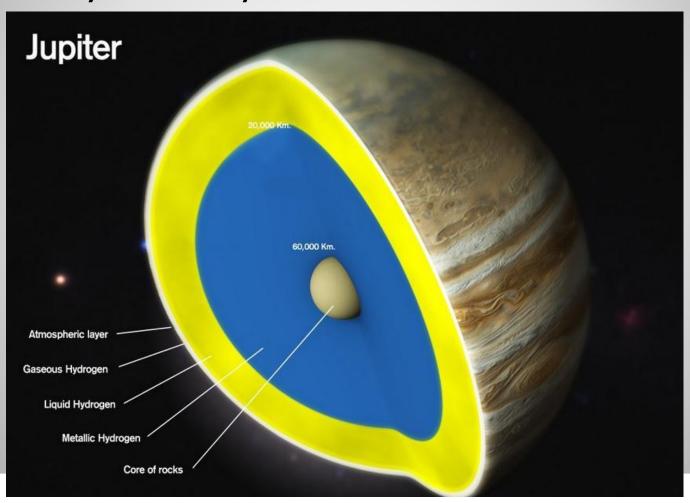
We figured this out using Earth's density,
 calculating the silicate and \_\_\_\_\_ densities and
 cross comparing them with the waves...

it worked!



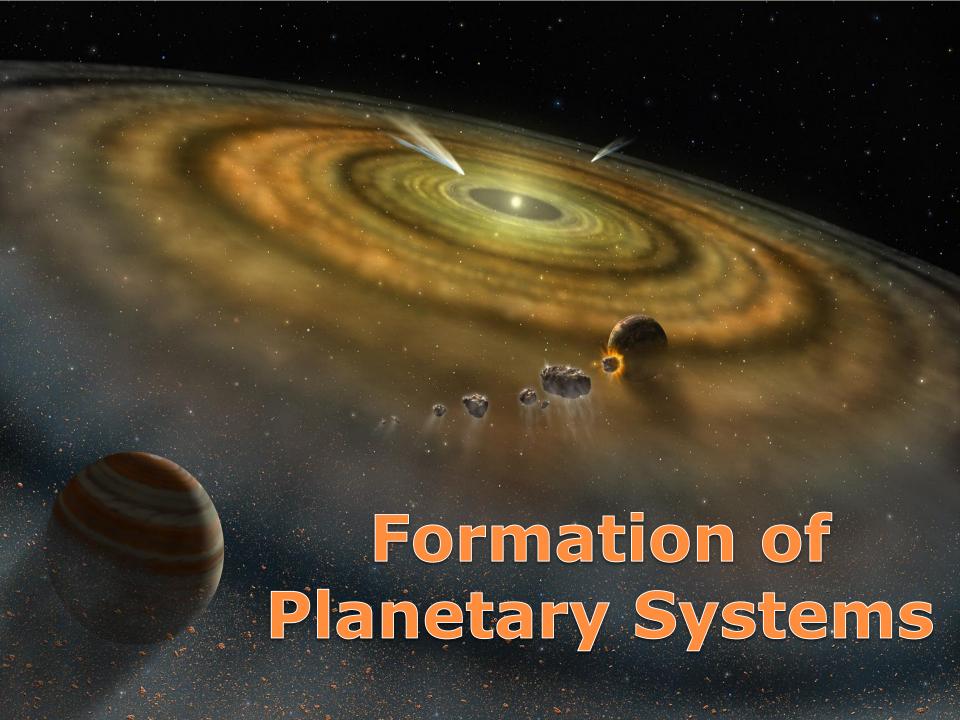
- of this strategy:
  - Several different \_\_\_\_\_ that will produce an equally good match to the observed density
  - The \_\_\_\_\_ of a given material can be affected by the planet's \_\_\_\_\_ force
- Conclusion:
  - All of the \_\_\_\_\_ planets have a similar density to \_\_\_\_\_ (about 3.9 to 5.5 g/cm³)
    - Largely rock with \_\_\_\_\_ core

- Conclusion
  - All of the \_\_\_\_\_ planets have a much smaller \_\_\_\_ (.7 to 1.7 g/cm<sup>3</sup>) similar to ice
    - Contain mainly methane, ammonia, and ice (H<sub>2</sub>O)
  - Probably have an \_\_\_\_\_ core and \_\_\_\_\_
     base the size of Earth on the inside
    - This was figured out based on the mass calculations and the effects of gravitational pull that these planets can create
    - Jupiter estimated core \_\_\_\_\_ the mass of Earth... something solid has to be there...



Age of the Solar System

- Outside of their differences in \_\_\_\_\_\_\_\_, and \_\_\_\_\_\_\_, it seems as though almost everything in the solar system formed at nearly the same time
- We can directly measure that date for the Earth, \_\_\_\_\_\_, and some asteroids
  - Thanks to \_\_\_\_\_ of their rocks
  - None are more than 4.6 billion years old
- The \_\_\_\_\_ is our age, too
  - Based off of its current brightness, temp., and rate of nuclear fuel consumption



## Formation of the Solar Systems

Introduction

• easy to figure out			
<ul><li>Why? We weren't there to it</li></ul>			
<ul> <li>Whatever we come up with has to support</li> </ul>			
these of the solar system:			
<ul> <li>The system is, with all of the planets orbiting the same</li> </ul>			
• There are two types of planets, and			
<ul> <li>With the inner being rock and outer being ice and gas</li> </ul>			
<ul> <li>The of the outer planets isn't too far off of</li> </ul>			
the, and the same is true for the inners			
(minus the gas)			
<ul> <li>All of the whose ages have so far been</li> </ul>			
determined are younger than 4.6x109 old			

## Formation of the Solar Systems

Introduction

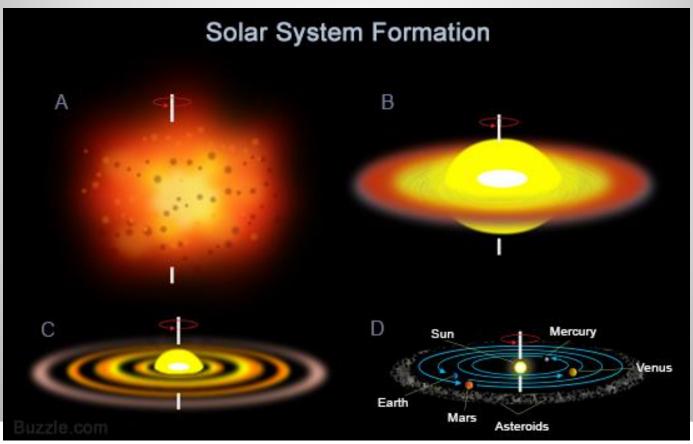
<ul> <li>Top theory:</li> </ul>	Theory
---------------------------------	--------

- States that the \_\_\_\_\_ system originated from a \_\_\_\_\_, flattened \_\_\_\_\_ of gas and dust, with the outer part of the disk becoming the planets and the center becoming the \_\_\_\_\_
  - Supports: the horizontal plane and the counterclockwise orbit of all of the planets
- We \_\_\_\_\_ that if there are other solar systems out there that could be similar to ours, their properties must be \_\_\_\_\_

## Formation of the Solar Systems

Introduction

Top theory: Solar Nebula Theory



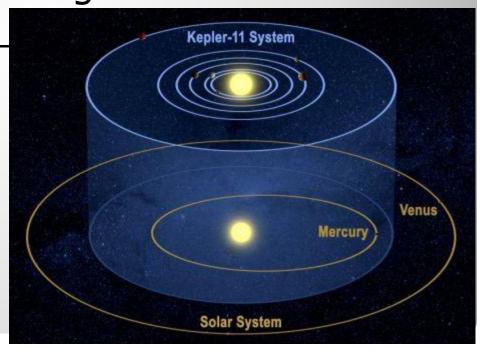
Introduction

Top theory: Solar Nebula Theory

We are searching for and \_\_\_\_\_ other

in various stages to see if these

stages are \_\_\_\_\_



Interstellar Clouds

- Interstellar Cloud \_\_\_\_\_ rotating
   \_\_\_\_\_ (whole combo) of \_\_\_\_\_ and dust
  - Common between stars and astronomers believe these are what developed into each of the stars
  - Right now, \_\_\_\_\_ stars could have planets orbiting them... we have no way to know for sure right now
    - Both the stars and the planets would have developed from that dust and gas

Interstellar Clouds

- The cloud that developed into the \_\_\_\_\_\_ was probably every bit of a couple light years in \_\_\_\_\_ and twice the present mass of the \_\_\_\_\_
- Interstellar Grains tiny dust particles found amongst the \_\_\_\_\_ in interstellar clouds
  - Combo of: \_\_\_\_\_\_, iron, \_\_\_\_\_, and frozen
  - These elements have been shown in the same proportions of the Sun according to the Sun's

Interstellar Clouds

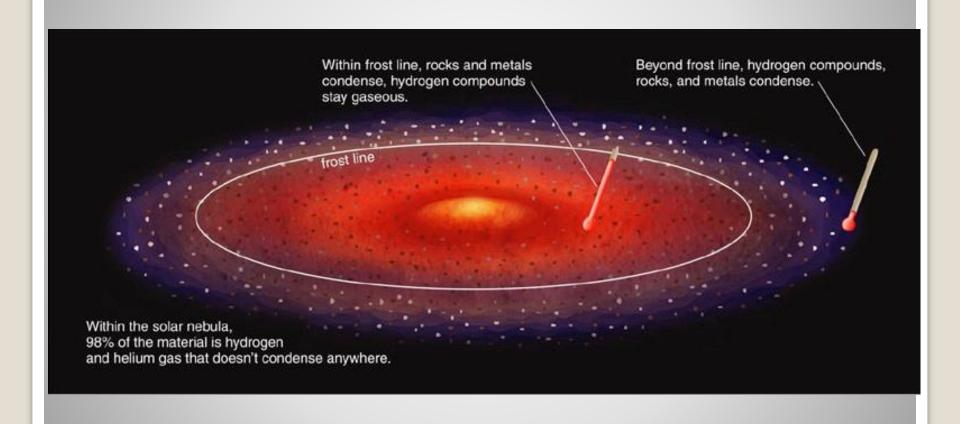
- The cloud began \_\_\_\_\_ into the Sun and planets when the gravitational attraction between the \_\_\_\_\_ in the \_\_\_\_ parts of the cloud caused it to collapse inward
  - Could've been triggered by a star \_\_\_\_\_\_
     nearby or hitting into another cloud
  - Because the cloud was \_\_\_\_\_\_, it became flat rather than fully collapsing in the middle



Formation of the Solar Nebula

- Solar Nebula rotating \_\_\_\_\_ with a \_\_\_\_ at the center from a collapsing cloud
   Took a few \_\_\_\_ years to occur
  - Condensed into the planets while the bulge became the sun
  - This supports the disk-like structure and the orbit pattern of the planets
  - Probably about \_\_\_\_\_ AU in diameter and \_\_\_\_\_
     10 AU thick
  - Some areas were \_\_\_\_\_ hot (especially the center) while others were well below the freezing point
  - We have been able to figure this out thanks to the Hubble and seeing the same set-up with other stars

#### Formation of the Solar Nebula



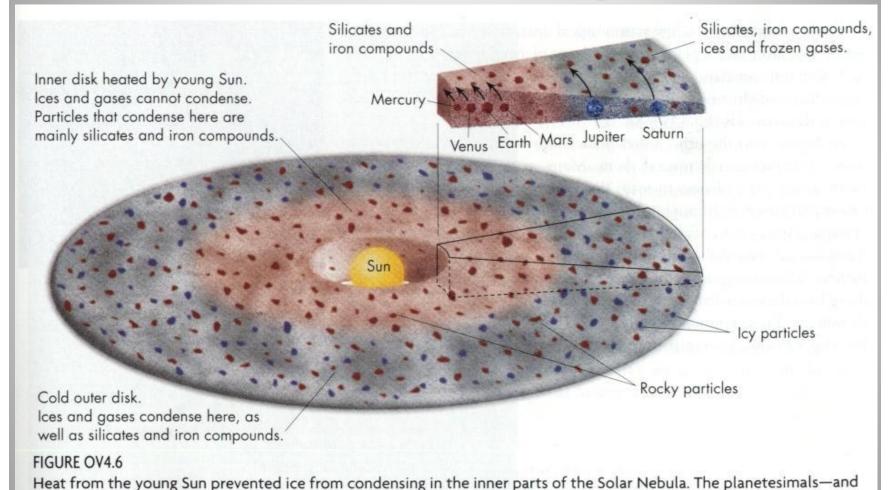
Condensation in the Solar Nebula

- Condensation occurs when a gas \_\_\_\_\_ and its molecules stick \_\_\_\_\_ to form a liquid or solid
  - There was an entire \_\_\_\_\_ sequence in the solar nebula as it cooled after collapsing
  - The \_\_\_\_\_ heat could only reach so far and that division in condensation created the inner and outer planets
    - The silicate-iron particles in the inner part
    - Similar \_\_\_\_\_ part but with \_\_\_\_\_

Accretion and Planetesimals

- Accretion when tiny \_\_\_\_\_ that condensed from the must have begun to stick \_\_\_\_\_ into bigger pieces This eventually created the planetesimals Planetesimals – small \_\_\_\_\_ bodies Perhaps held together by electrical forces like static electricity (that weren't too crazy) allowed particles to stick together, too
  - Range in size from a few \_\_\_\_\_ to km

ultimately the planets—that formed there are therefore composed mainly of rock and iron.



Formation of the Planets

•	officiation of the flattets
	As the planetesimals moved within the
	and, planets began to form
	<ul> <li>Some hit and while others that collided more gently stuck together</li> </ul>
	<ul> <li>Due to gravity, substances found in certain areas of the and other chemical factors (like</li> </ul>
	) it has been concluded that these collisions are what lead the outer planets to be so much larger than the ones
	<ul> <li>Almost all of the planets like Earth, but the inner planets couldn't hold the gas layers like the outers and therefore they became much larger in volume</li> </ul>

Formation of Satellite Systems

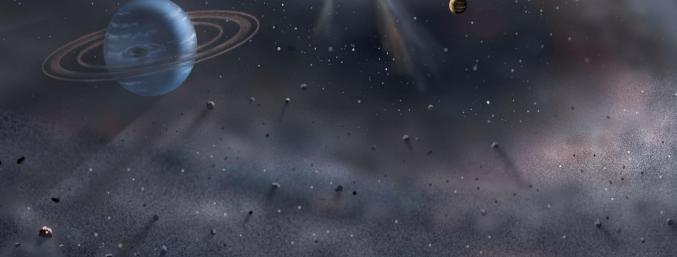
- The satellite systems include the \_\_\_\_\_
   and other materials that orbit
  - This developed once the \_\_\_\_\_ was able to develop a larger \_\_\_\_\_, strong enough to begin attracting other objects to itself
  - Many of the satellites (moons) are about as large as Mercury and would be considered \_\_\_\_\_\_ if they orbited the sun rather than another planet



Formation of the Atmospheres

- Last part of the \_\_\_\_\_
- and outer planets are thought to have formed \_\_\_\_\_ differently
  - Outer: captured the \_\_\_\_\_ from the nebula
  - Inner: not \_\_\_\_\_ enough and too hot to capture the gas from the nebula
    - Likely created their own from \_\_\_\_\_ and by retaining gases from \_\_\_\_\_ comets and icy planetesimals that vaporized on contact





#### Introduction

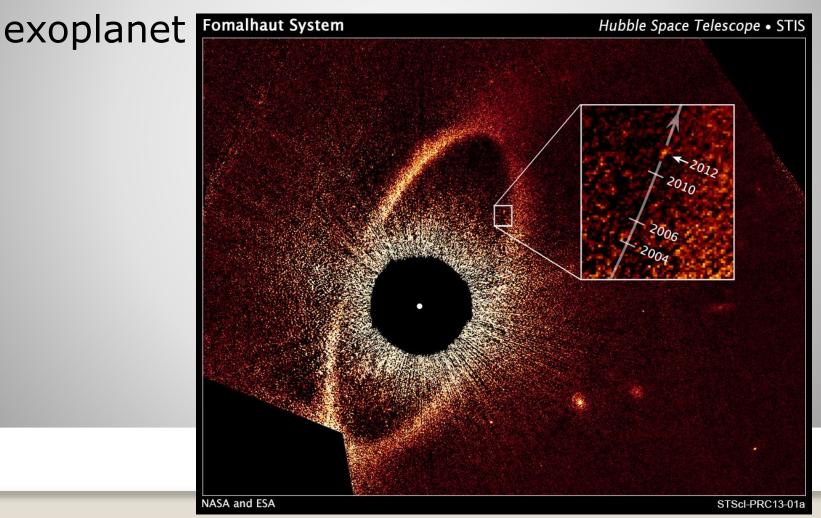
- Exoplanets \_\_\_\_\_ orbiting \_\_\_\_\_
   other than the Sun
  - Studying other planets helps us better understand how our solar system \_\_\_\_\_
  - Most present \_\_\_\_\_ for exoplanets comes from their effect on the star they orbit
    - The planet exerts a gravitational force back on the star as a result of Newton's third law (actionreaction)
    - This causes the star to wobble which creates a
       \_\_\_\_\_ shift in the spectra that we can measure

#### Introduction

- Fomalhaut a star with a \_\_\_\_\_
   exoplanet
  - Estimated to be a \_\_\_\_\_ million year old star
  - Has an ice \_\_\_\_\_ around it (similar to Kuiper's belt)
  - The planet is assumed to have been \_\_\_\_\_ by accreting that frozen material
  - The exoplanet is really faint and hard to see against the star's light, but the evidence from the star, itself, tells us a lot about that exoplanet

#### Introduction

Fomalhaut – a star with a detected

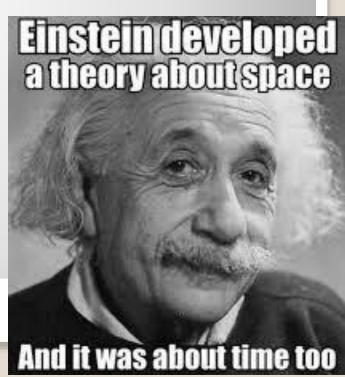


Finding Exoplanets

- Many of the objects we pick up on are
   in size and close to their
- This allows us to pick up on that Doppler signal since most are hardly \_\_\_\_\_, even being large in size.

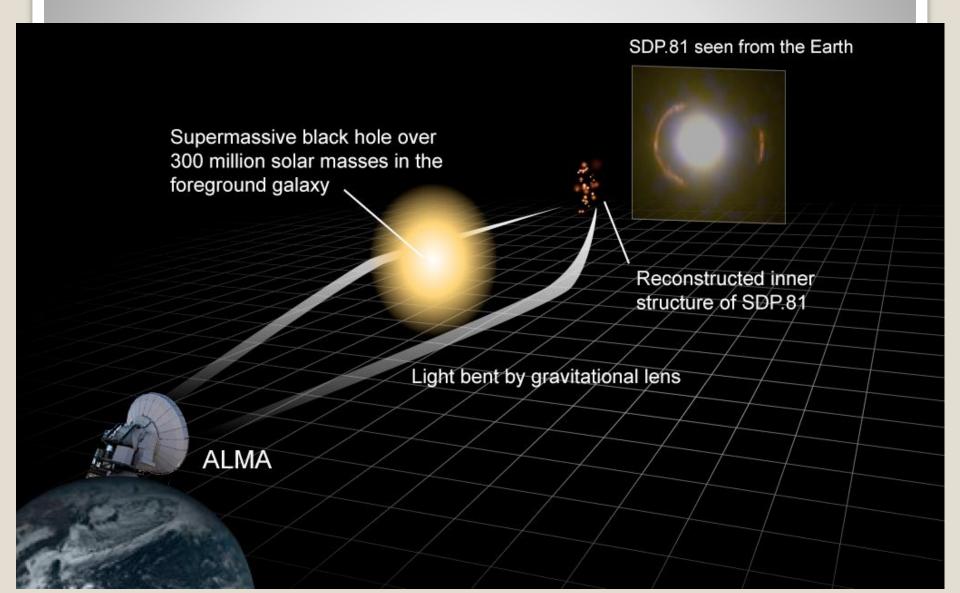
being large in size

Another idea...use \_\_\_\_\_\_ approach



Einstein's Approach

- He showed that a mass \_\_\_\_\_ space in its \_\_\_\_\_ and that this bending creates the mass's gravity
  - Part of his general theory of \_\_\_\_\_
- If a \_\_\_\_\_ of light passes near a mass, the bent space around the mass deflects the light and can bring it to a \_\_\_\_\_



Einstein's Approach

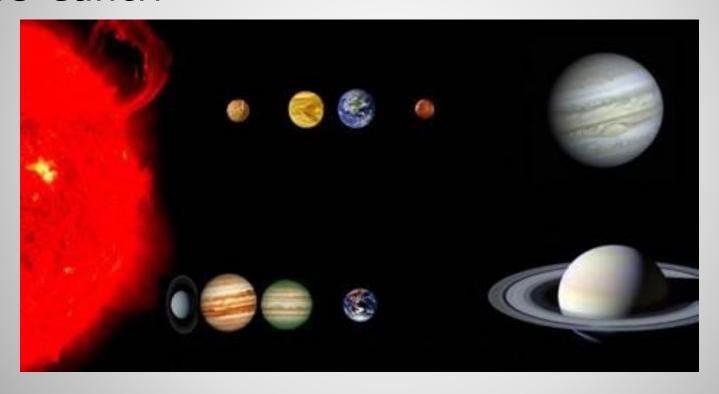
- Gravitational Lensing \_\_\_\_\_ of light by gravity
  - Great tool for detecting \_\_\_\_\_ planets
  - How it works:
    - Measure a star's \_\_\_\_\_
    - If a planet crosses in between, its mass will bend the light and because of reflection actually focus more light our way (it's not much more, but hey, any little bit helps)
- Astronomers are running \_\_\_\_\_ of data screens on \_\_\_\_ of stars to detect any slight \_\_\_\_ in that brightness level that might suggest a planet or big body that would be present

**Exoplanet Systems** 

- As of what has been released to the
   \_\_\_\_\_, astronomers haven't found a
   system that looks \_\_\_\_\_\_ like our own
- The nearest match so far is the system of planets orbiting the star 55
  - This sun-like star has five planets orbiting within
     AU of the star just like ours
  - All of the planets are massive (10x Earth)
  - 3 of the planets orbit closer to the star than does from the Sun

**Exoplanet Systems** 

55 Cancri



#### **Exoplanet Systems**

- 55 Cancri and its set up really \_\_\_\_\_ our understanding of how the solar system set itself up
  - having \_\_\_\_\_ gas planets so close to the star
  - According to solar nebula, they should've formed much \_\_\_\_\_ back off of the star where the temperatures are much lower
  - Astronomers are working on understanding what's different in \_\_\_\_\_ like that one that make this scenario possible in its solar system and not ours

**Exoplanet Systems** 

- It is thought that planets in other systems might have the ability to "\_\_\_\_\_\_ " within the system
- Others are known to have \_\_\_\_\_ orbits rather than circular ones which can damage or effect the orbits of smaller planets in that same
  - This can either eject them out or cause them to crash into the star

The \_\_\_\_\_ continues....